Combinations of Earth orientation Measurements: SPACE95, COMB95, and POLE95

R S Gross (Jet Propulsion 1 laboratory, California Institute of Technology, 4800 Oak Grove Drive, Pasadena, CA 91 109; ph. 818-354-4010; fax:818-393-6890; email:rsg@logos.jpl.nasa.gov)

A Kalman filter has been used to combine independently determined measurements of the Earth's orientation take. by the modern, spacegeodetic techniques of very long baseline interferometry, satellite laser ranging, lunar laser ranging, and Global Positioning System interferometry. Prior to combining the measurements, UT1 tidal t erms were removed, out ying dat a points were deleted, series-specific corrections were applied for bias and rate, and the stated uncertainties of the measurements were adjusted by multiplying them by seriesspecific scale factors. Values for these bias-rate corrections and uncertainty scale factors were determined by an iterative, round-robin procedure wherein each data set is compared, in turn, to a combination of all other data sets. When applied to the measurements, the bias-rate. corrections thus determined make the data sets agree with each other in bias and rate, and the uncertainty scale factors thus determined make the residual of each series (when differenced with a combination of ail others) have a reduced chi-square of one. The corrected and adjusted series are then placed within an IERS reference frame by aligning them with the IERS Earth orientation series 1 X)1'(11 ERS) C 04, The resulting combination of corrected, adjusted, and aligned Earth orientation series is designated SPACE95 and spans October 6,(), 1976 to February 1().(), 1996 al daily intervals.

Two additional combinations of Earth orientation measurements have been generated by combining optical astrometric measurements with the space-geodetic measurements used to form SPACE95, COMB95, formed by additionally incorporating the BHI optical astrometric measurements, spans January?(), (), 1962 to February 6.(), 1996 at5-day intervals. POLE95, formed by additionally incorporating both the BHI land II, S optical astrometric measurements, consists of just the polar motion tom]-mncmts of the 1 farth's orientation and spans 1900-" 1995 at monthly intervals.